

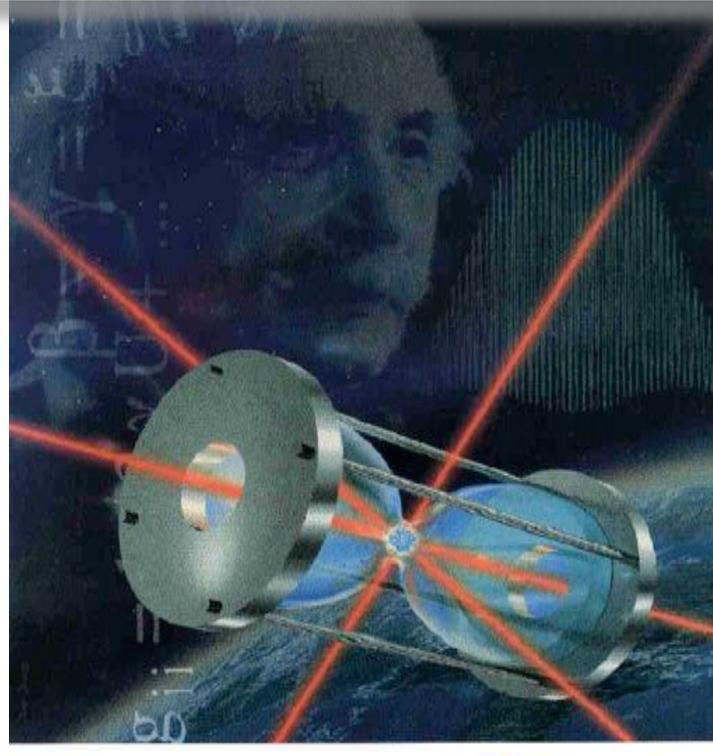
ACES: Ultra-stable Clocks in Space

Fundamental physics and Applications

1997



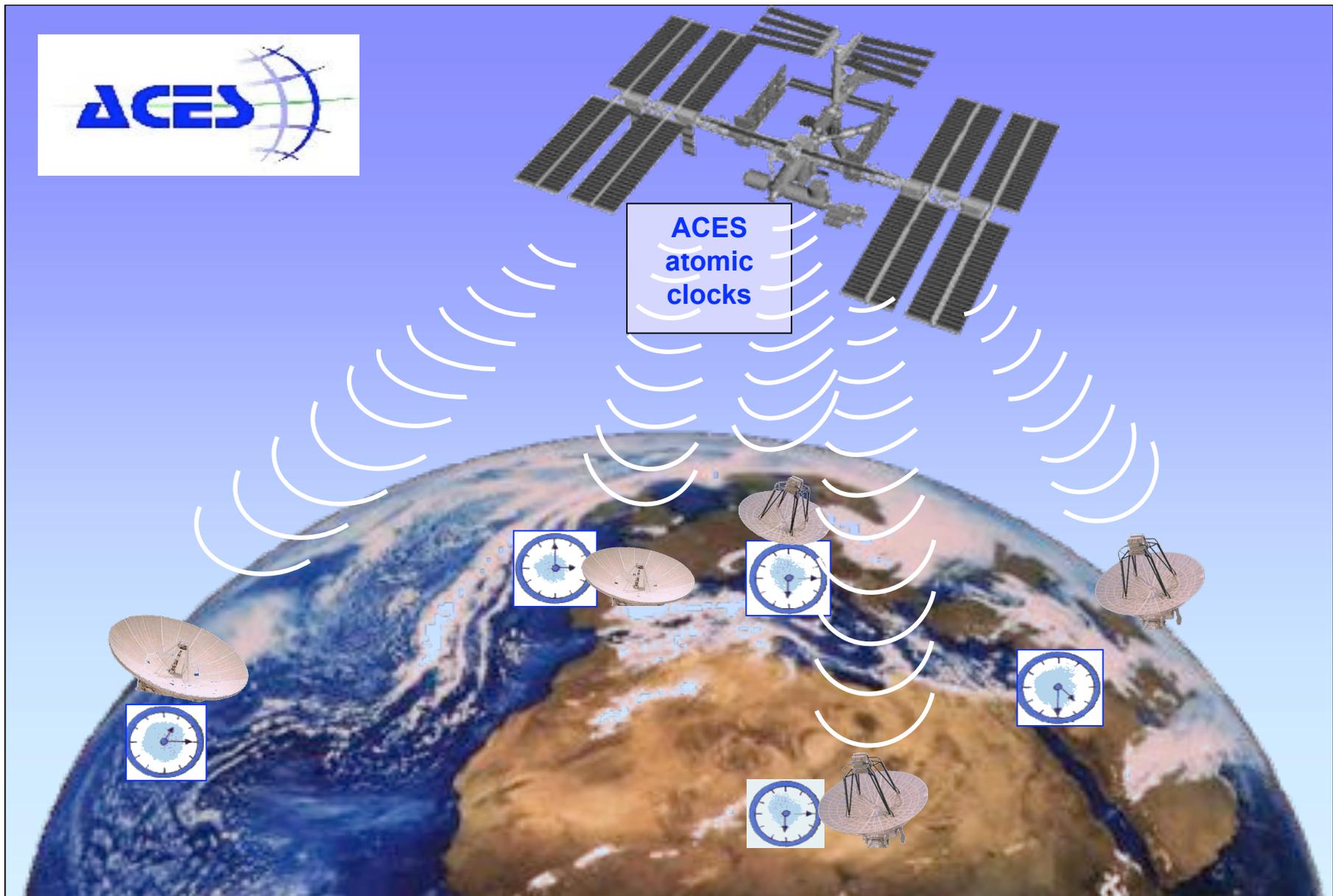
IGS, May 11th
2006



CENTRE NATIONAL D'ÉTUDES SPATIALES



C. Salomon, Ecole Normale Supérieure, Paris, France



- A cold atom Cs standard in space
- Fundamental physics tests
- Worldwide access



ACES instruments



The ACES experiment consists in 2 instruments plus “tools”:

1. **Pharao: Cold atom clock**

⇒ Laser cooled cesium clock designed for micro-gravity



2. **Space Hydrogen Maser – SHM**

⇒ Reference clock and local oscillator for Pharao

ON

3. **Frequency Comparison and Distribution Package – FCDP**

⇒ Frequency comparison and processing



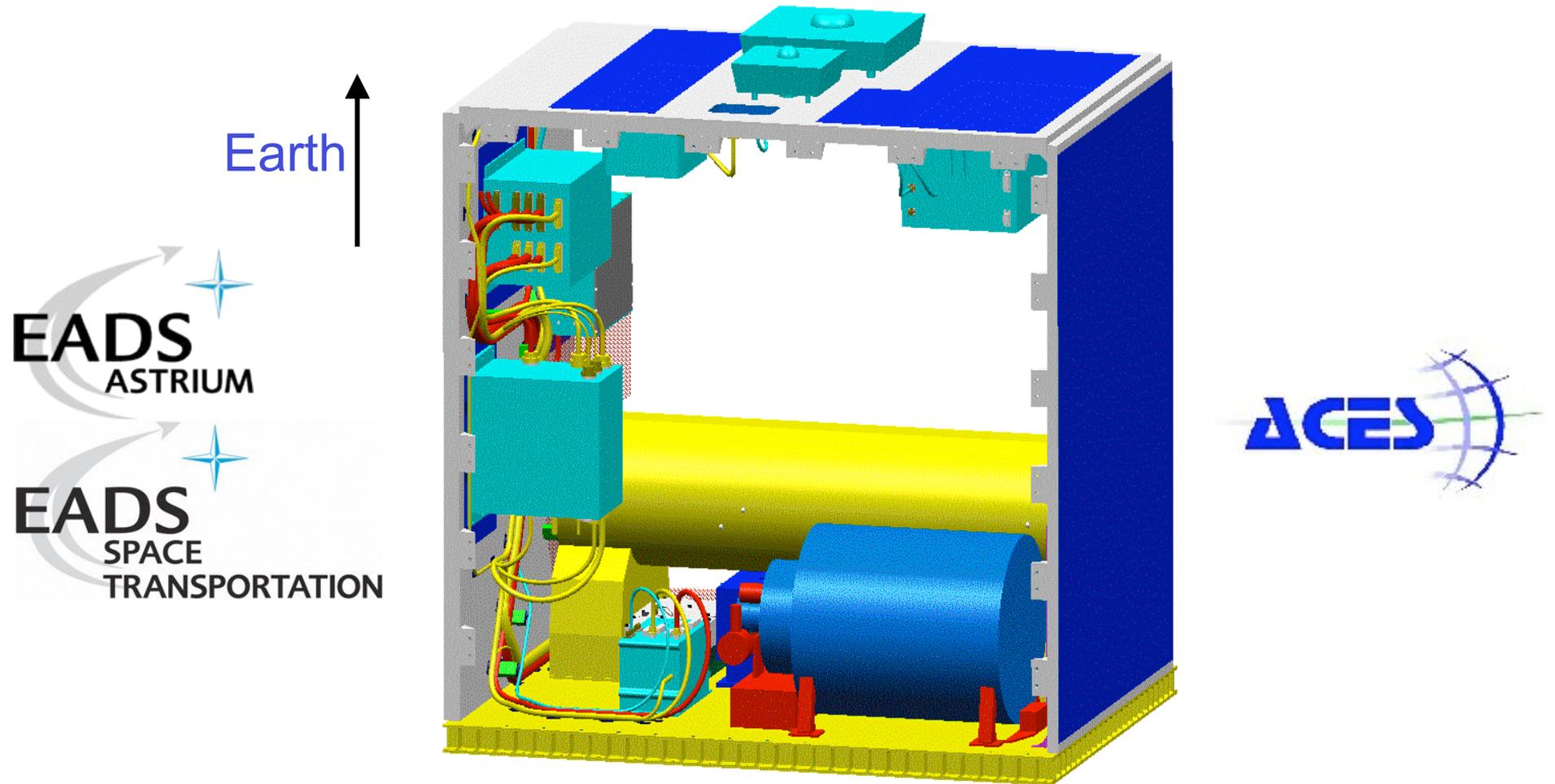
4. **Micro-Wave Link – MWL**

⇒ Link for time-frequency transfer to the ground

W. Schäfer talk



ACES General View



$M = 227 \text{ kg}$

$P = 450 \text{ W}$

ACES ON COLUMBUS EXTERNAL PLATFORM



ACES

Current launch date : 2010
Mission duration : 18 months



ACES OBJECTIVES (1)

1. Operate SHM and a cold atom clock in microgravity : PHARAO

- A linewidth of 100 millihertz
- A frequency stability of : $\sigma_y(\tau) = 7 \cdot 10^{-14}$ to $1 \cdot 10^{-13} \tau^{-1/2}$
 $< 3 \cdot 10^{-16}/\text{day}$

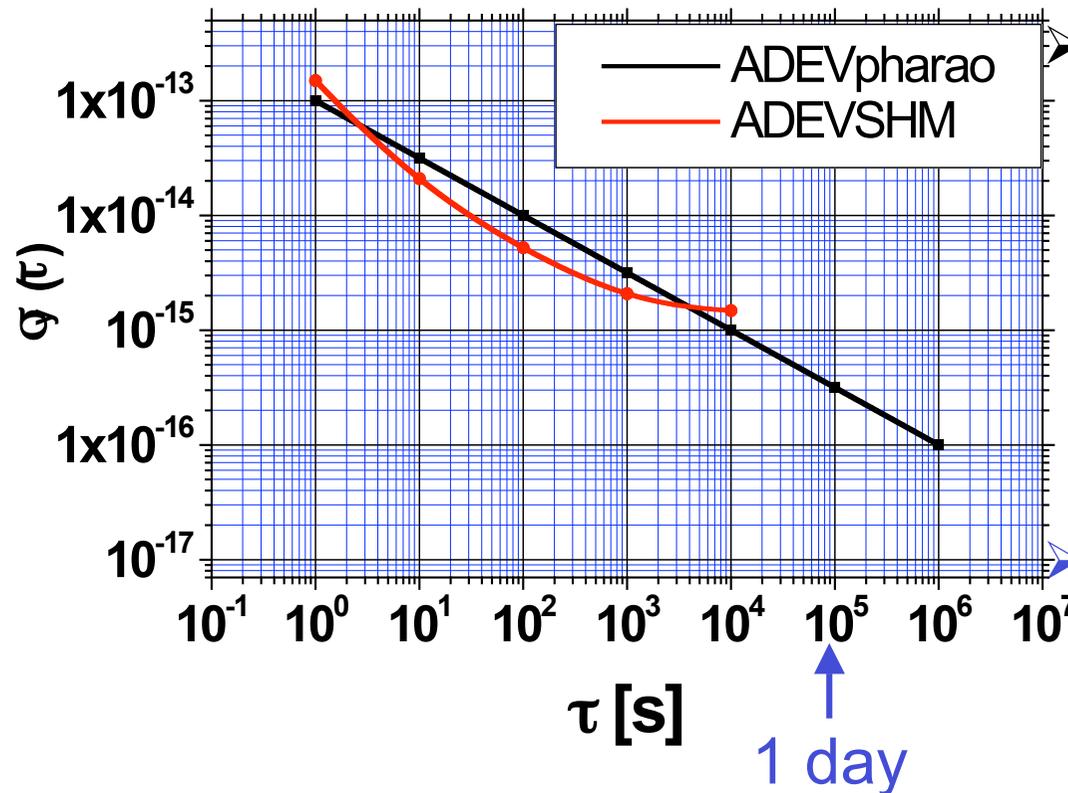
2. Study the ultimate stability and accuracy in space :

- Accuracy : $\sim 10^{-16}$

PHARAO performance is established through onboard comparison with SHM in FCDP and with ground clocks using MWL

Frequency stability of ACES Clocks

Allan deviation of the 2 clocks:



Stability of SHM in time intervals of 3 to 3000 s, well adapted to:

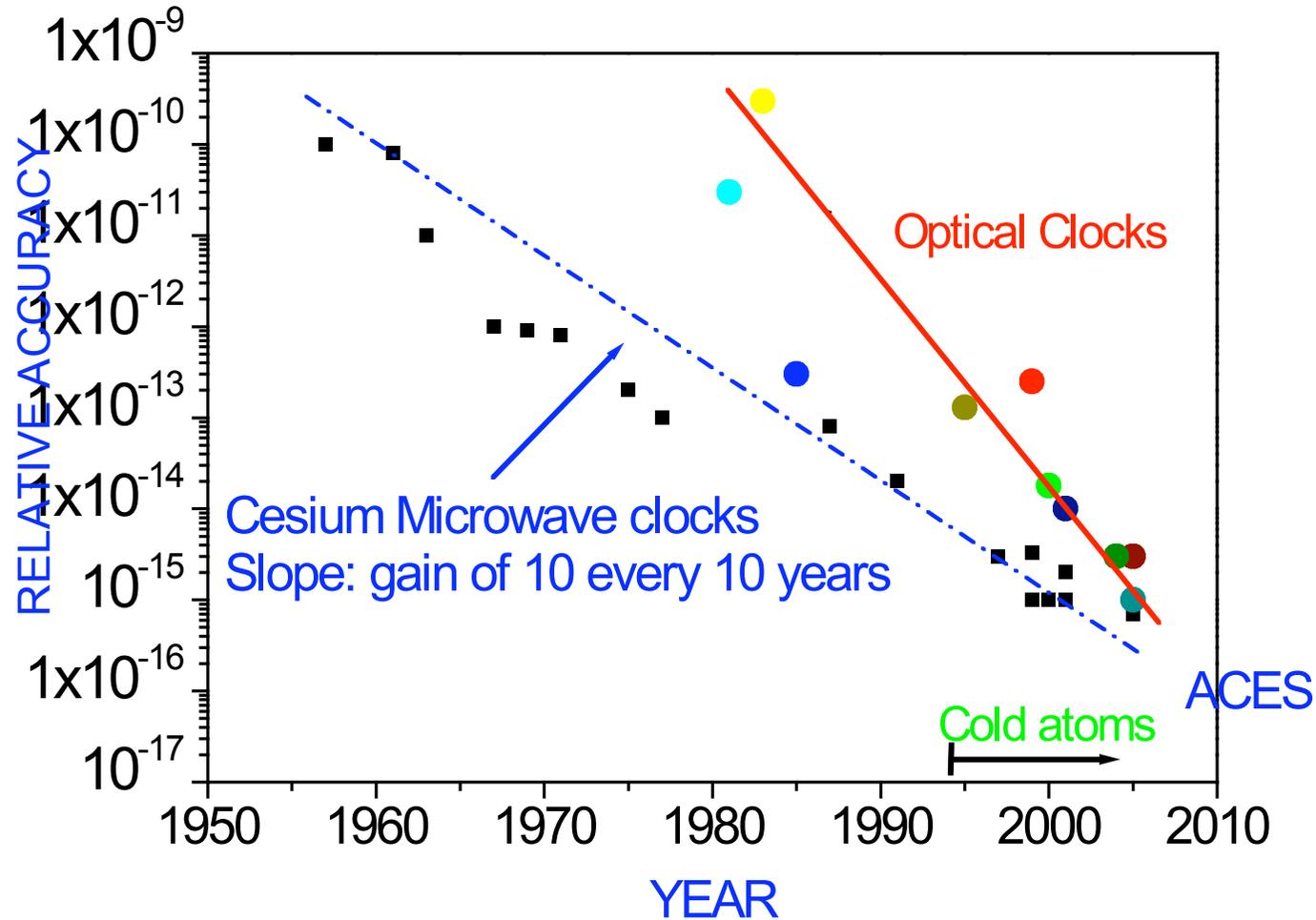
- ISS single pass (duration: 200-400 s)
- evaluation and optimization of PHARAO

ACES: slow servo of SHM onto PHARAO

Stability at one day: 3×10^{-16}
at 10 days: 1×10^{-16}

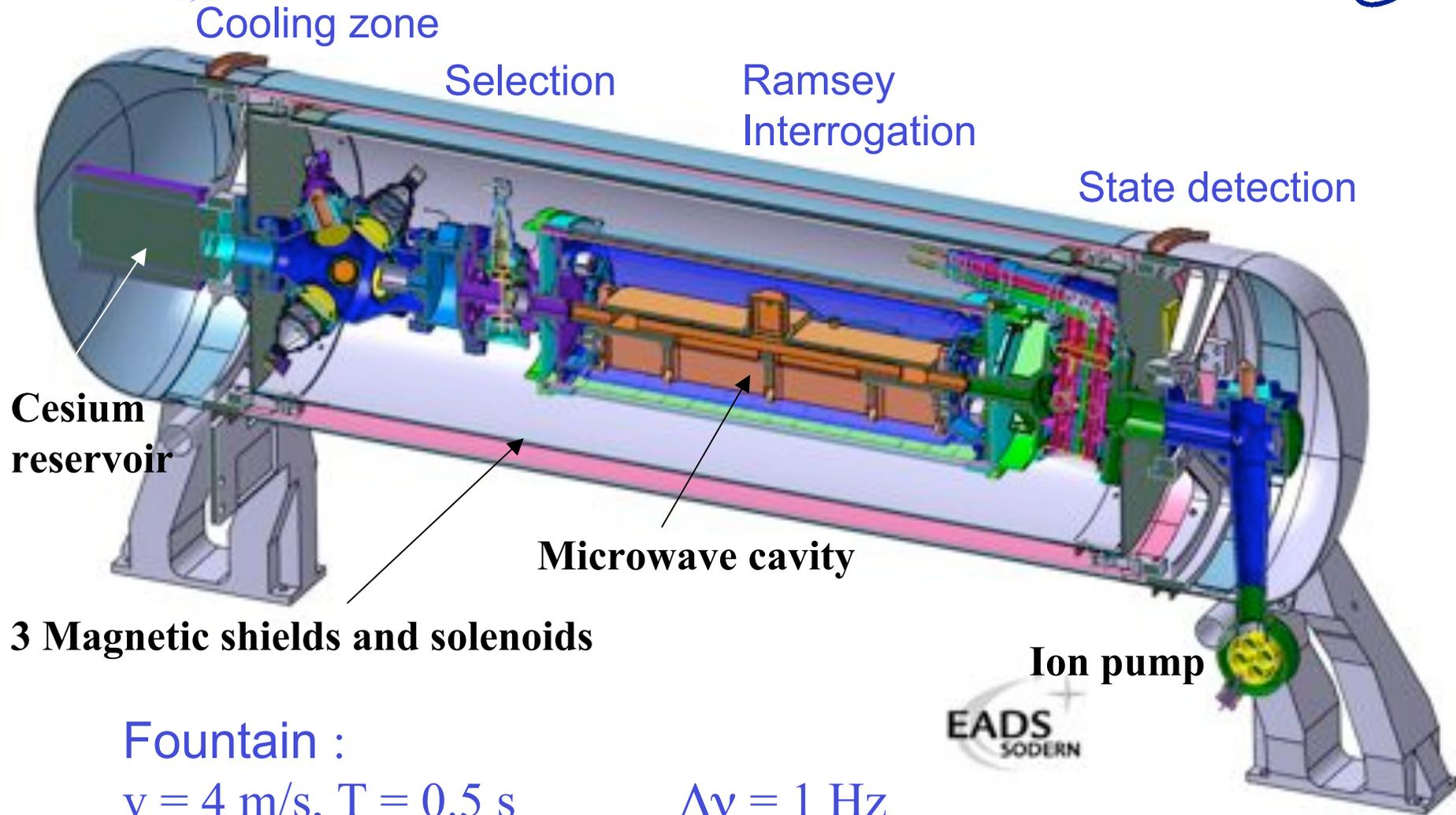
Accuracy of the atomic time

ACCURACY OF THE ATOMIC TIME



Current accuracy: 7×10^{-16}





Fountain :

$$v = 4 \text{ m/s}, T = 0.5 \text{ s}$$

$$\Delta\nu = 1 \text{ Hz}$$

- PHARAO :

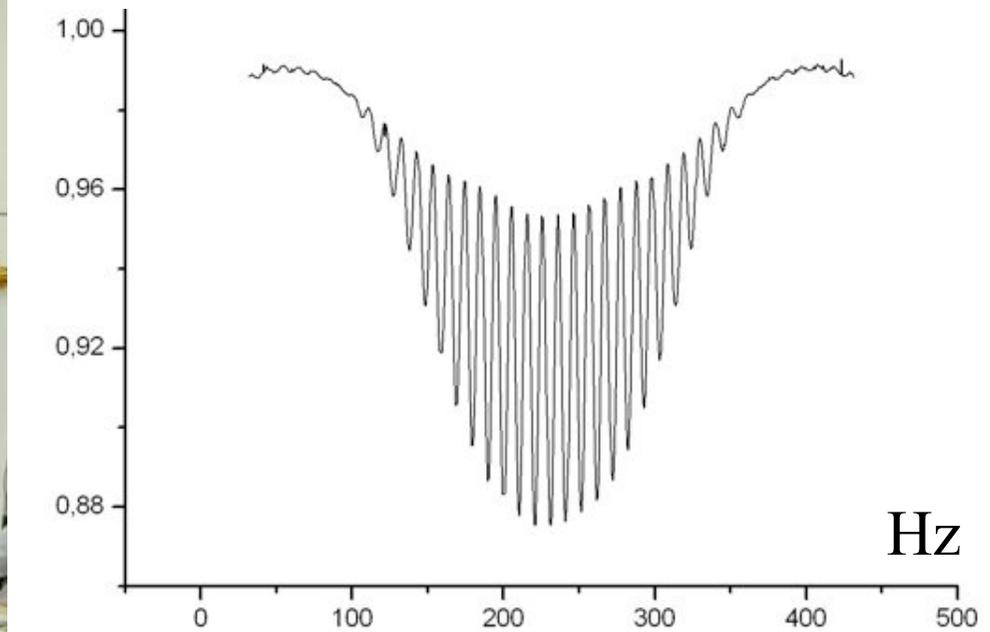
$$v = 0.05 \text{ m/s}, T = 5 \text{ s}$$

$$\Delta\nu = 0.1 \text{ Hz}$$

PHARAO Space Clock: first cold atoms !



Laser source



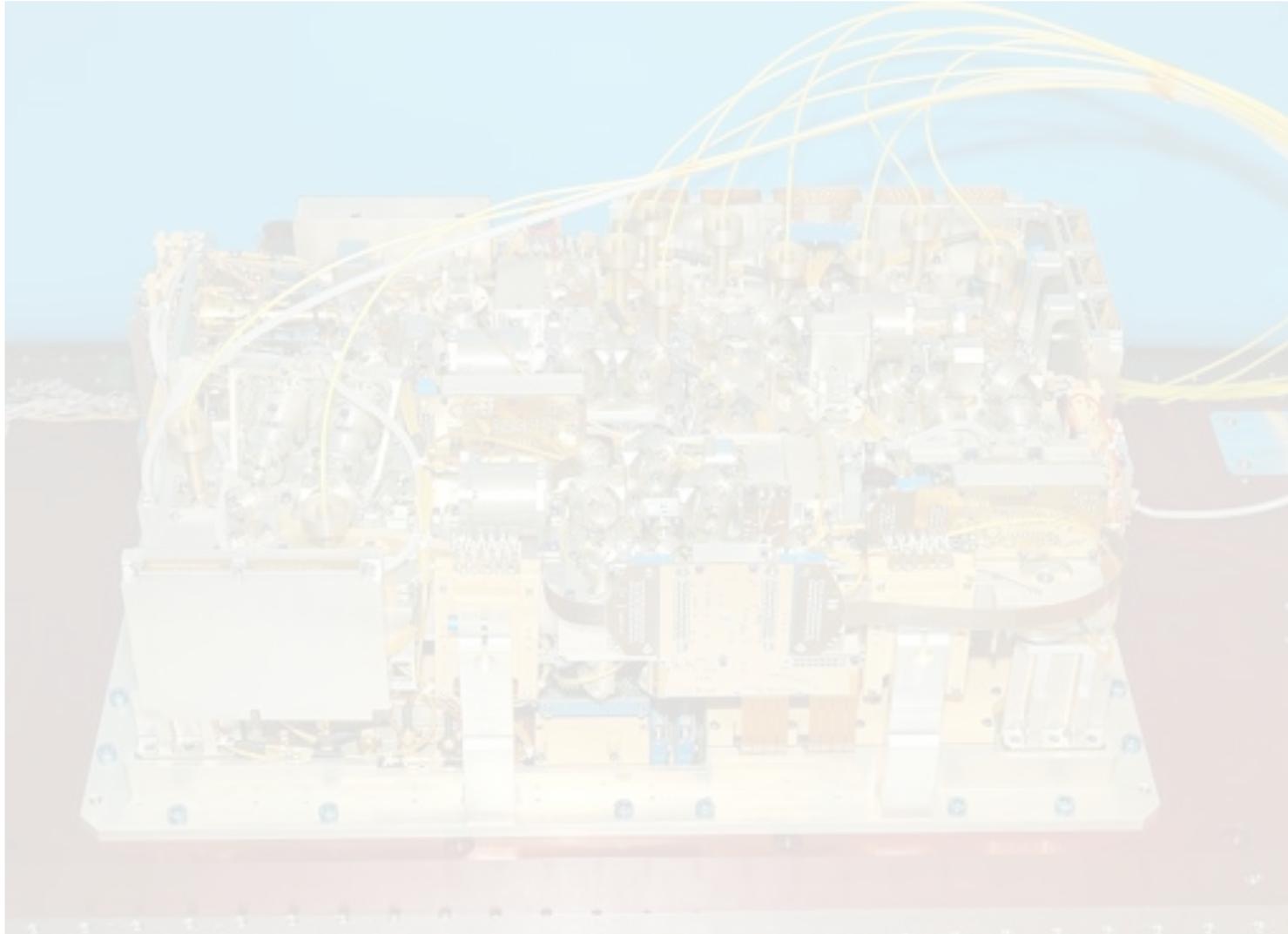
Clock Ramsey resonance

Cesium tube

Functional tests starting in CNES Toulouse

Laser Source

20.054 kg, 36W, 30 liters



Main active components:
4 ECDL
4 DL
6 AOM
30 PZT
11 motors
6 photodiodes
8 peltier coolers

ACES OBJECTIVES (2)

2. Ultra-stable frequency comparisons on a worldwide basis :

Clock comparisons@ 10^{-16}

Contribution to TAI

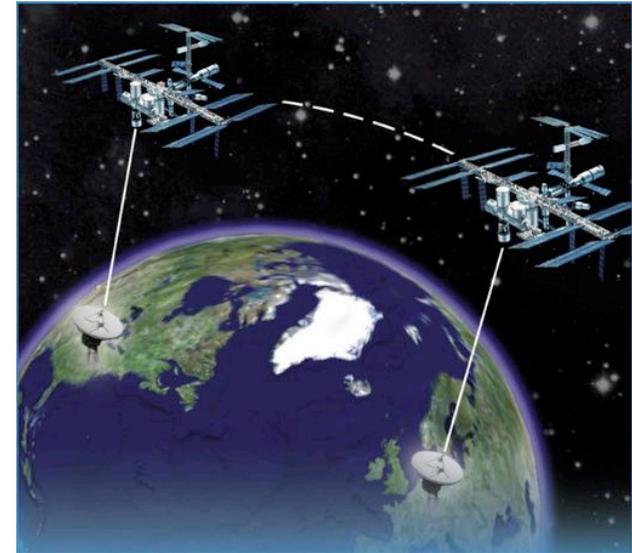
Space-ground and ground-ground comparisons

Gain: x 10 wrt current GPS.

Common view



non common view



ACES interface with GNSS and Geodesy

- Orbitography, [Svehla talk](#)
- Interest to install dual frequency GPS receiver connected to ACES [Montenbruck talk](#)
- Connection to IGS network
- Availability and interest of an ultra-stable reference clock in space
- MWL:Independent time transfer and comparisons with GPS/Galileo

ACES OBJECTIVES (3)

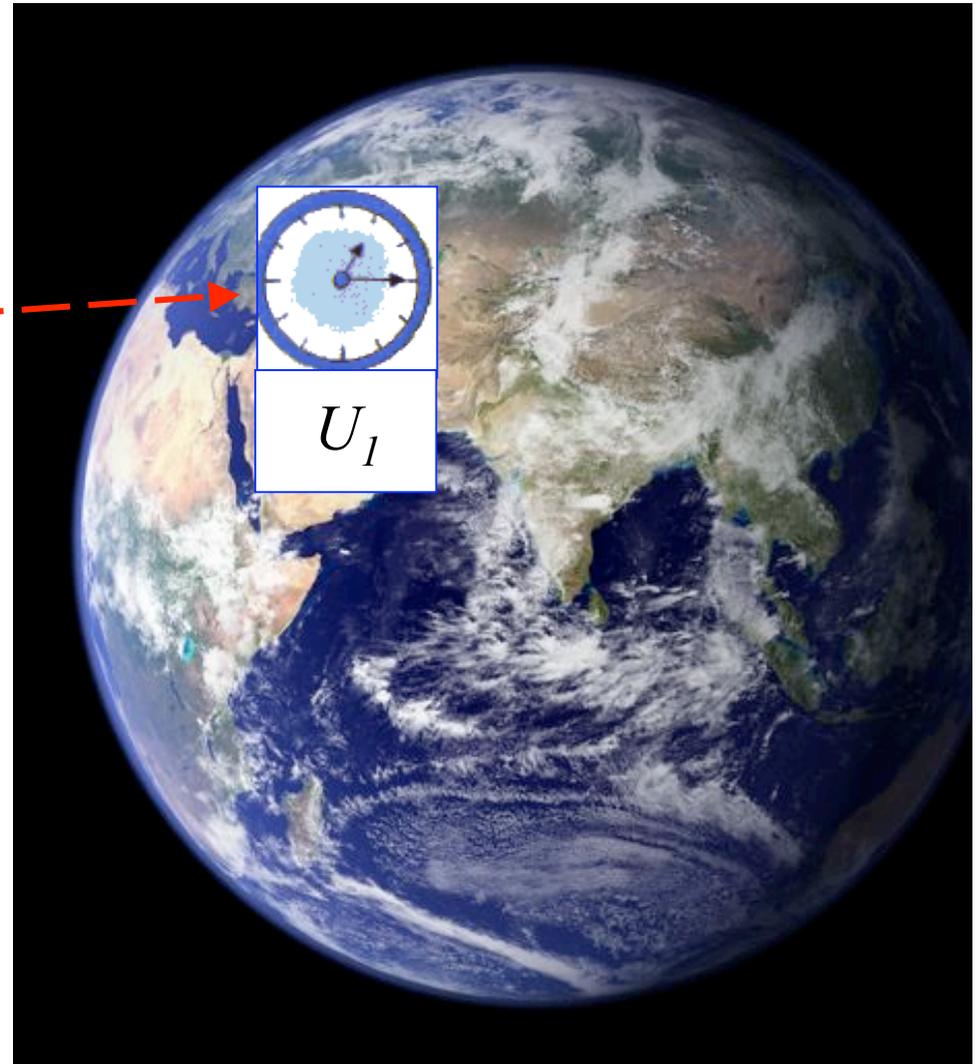
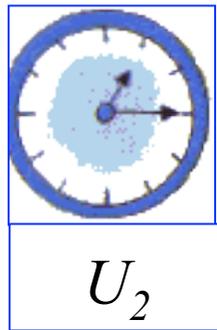
3. Test General Relativity :

Einstein effect: Red shift : x 35 sensitivity improvement

Search for a possible drift of the
fine structure constant α : 10^{-16} / year (x 10)

Test of Lorentz invariance (x10 to x 100)

A Prediction of General Relativity



$$\frac{i_2}{i_1} = \left(1 - \frac{U_2 - U_1}{c^2} \right)$$

Redshift measurement:
ACES target: $2 \cdot 10^{-6}$

Do fundamental physical constants vary with time ?

G , α_{elm} , α_{strong} , m_e, \dots

Principle : Compare two or several clocks of different nature as a function of time

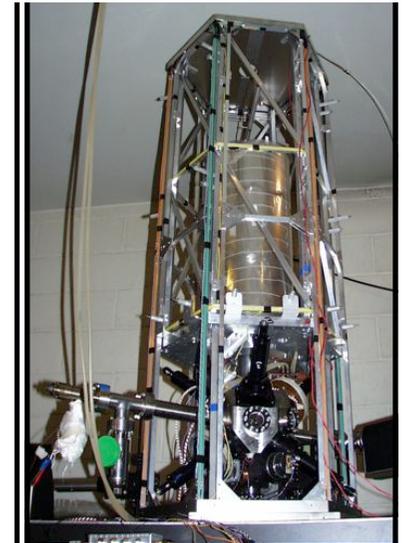
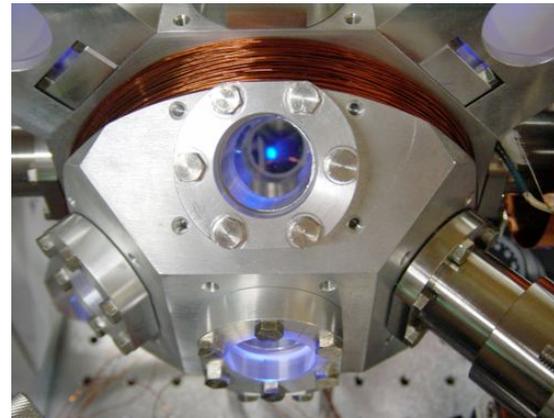
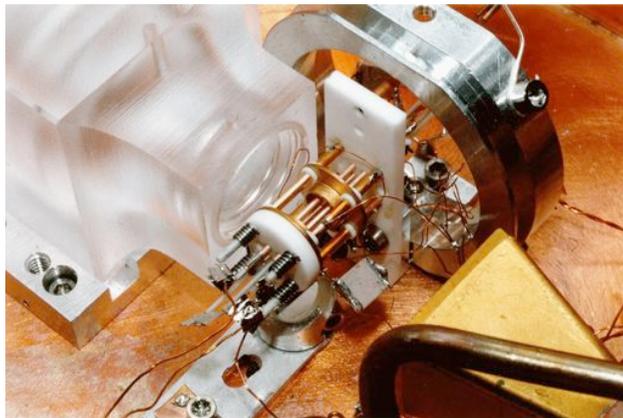
ex:

Microwave clock/Microwave clock

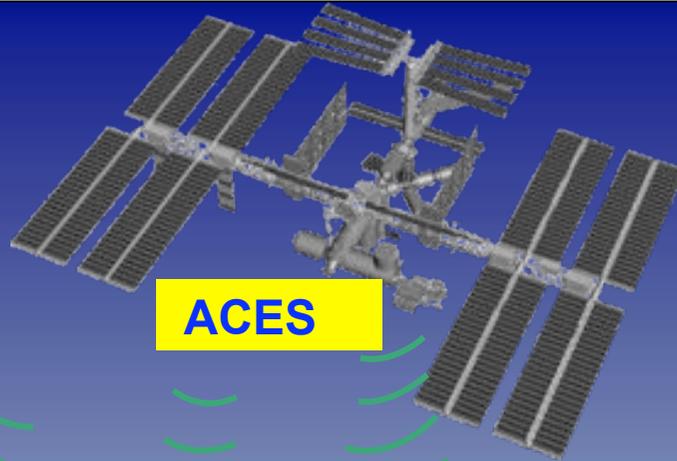
rubidium and cesium

Microwave/Optical clock

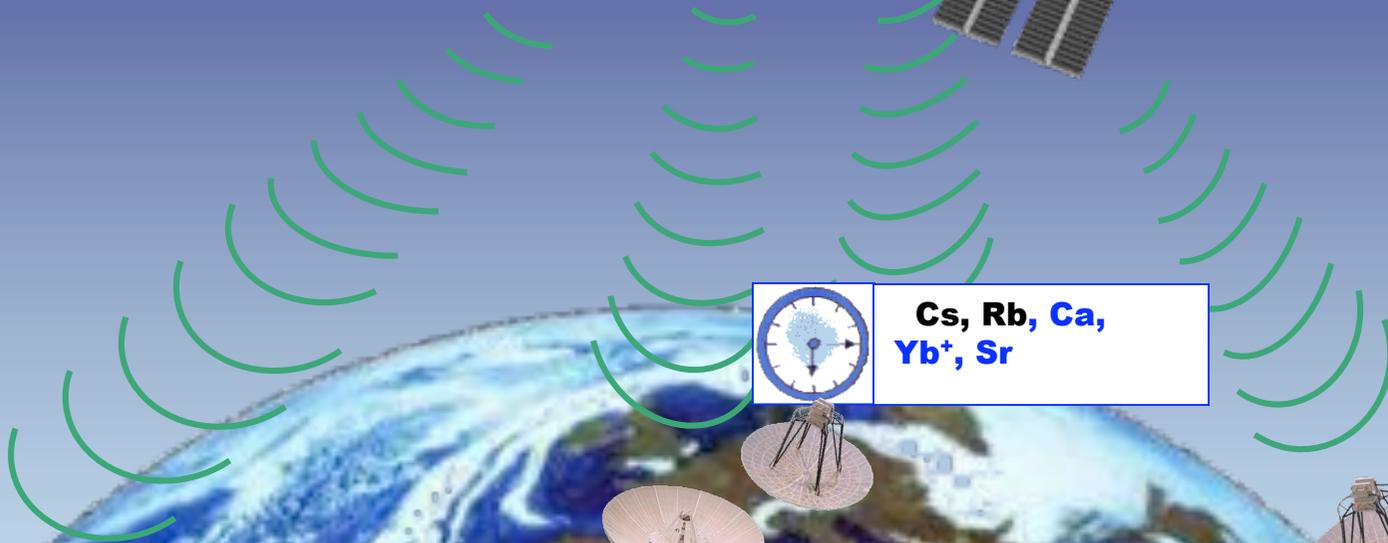
Optical Clock/Optical clock



The ovens and electrodes of the NPL strontium ion end-cap trap.



ACES



**Cs, Rb, Ca,
Yb⁺, Sr**



**Cs, Rb, Sr, Hg
H, In⁺, Mg, Ag**



Cs, Yb⁺, Yb⁺,



**Cs, Hg⁺
Al⁺, Sr,
Ca, Yb**



Cs, Rb



**Cs, Rb, Sr⁺,
Yb⁺**

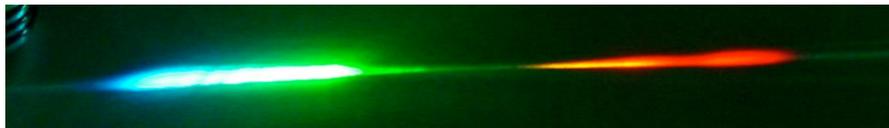


ACES : a link between different clocks at 10^{-16}

- Transition or oscillator
- solid resonator: R_y / α
- electronic transition: R_y
- fine structure transition: $\alpha^2 R_y$
- hyperfine transition: $g_p (m_e/m_p) \alpha^2 R_y$

With $R_y = \alpha^2 m_e c^2 / 2h$

Femtosecond laser system
for frequency comparisons



J. Hall
T. Hänsch

2005 Nobel prize for physics

ACES Ground Stations (May 06)

Australia:	UWA, CSIRO(Sydney)
Austria:	Univ. Innsbruck
Brazil:	Univ. Sao Carlos
Canada:	NRC
China:	Shangai Obs, NIM, NTSC
Germany:	PTB, MPQ, Univ. Hannover, Univ. Düsseldorf, TU Muenchen, Univ. Erlangen
France:	SYRTE, CNES, Obs. Besançon, OCA, LPL
Italy:	IEN, Univ. Firenze
Japan:	Tokyo Univ., NMIJ, CRL
Russia:	Vniftri, ILS Novosibirsk
Swiss:	METAS, ON
England:	NPL
USA:	JPL, NIST, Penn St. Univ., USNO
Taiwan:	Telecom research lab
Int. Agency:	BIPM

**Total : 34 institutes + theory groups
> 260 researchers**



Beyond ACES

Clocks on Earth at 10^{-17} will be limited by the knowledge of the local Earth potential and of its fluctuations.

Next step: dedicated satellites for global time dissemination

Fast advances in optical clocks: 10^{-17} becomes realistic

A new Relativistic Geodesy: based on red-shift

Space is a quiet environment: ultra-stable lasers, optical resonators and clocks (OPTIS)

Vastly improved tests of GR in solar orbit

SORT, Shapiro delay, differential redshift, ...

Space-Space VLBI with sub-micro arc-second resolution

New matter wave sensors

Bose Einstein Condensates in Space and Atom Lasers

